

# The Scintillation Prediction Observations Research Task (SPORT) Mission

James F. Spann (NASA/MSFC), Charles Swenson (USU), Otavio Durão (INPE), Luis Loures (ITA), Rod Heelis (UTD), Rebecca Bishop (Aerospace), Guan Le (NASA/GSFC), Abdu (INPE), Linda Krause (NASA/MSFC), Clezio Dinardini (INPE), Eloi Fonseca (ITA)



## Science

The Scintillation Prediction Observations Research Task (SPORT) mission tackles the very difficult problem of understanding the conditions under which ionospheric variability develops that leads to scintillation that compromises transmission signals. SPORT seeks to answer two science questions:

- What is the state of the ionosphere that gives rise to the growth of plasma irregularities that extend into and above the F-peak?
- How do plasma irregularities evolve to impact the appearance of radio scintillation at different frequencies?

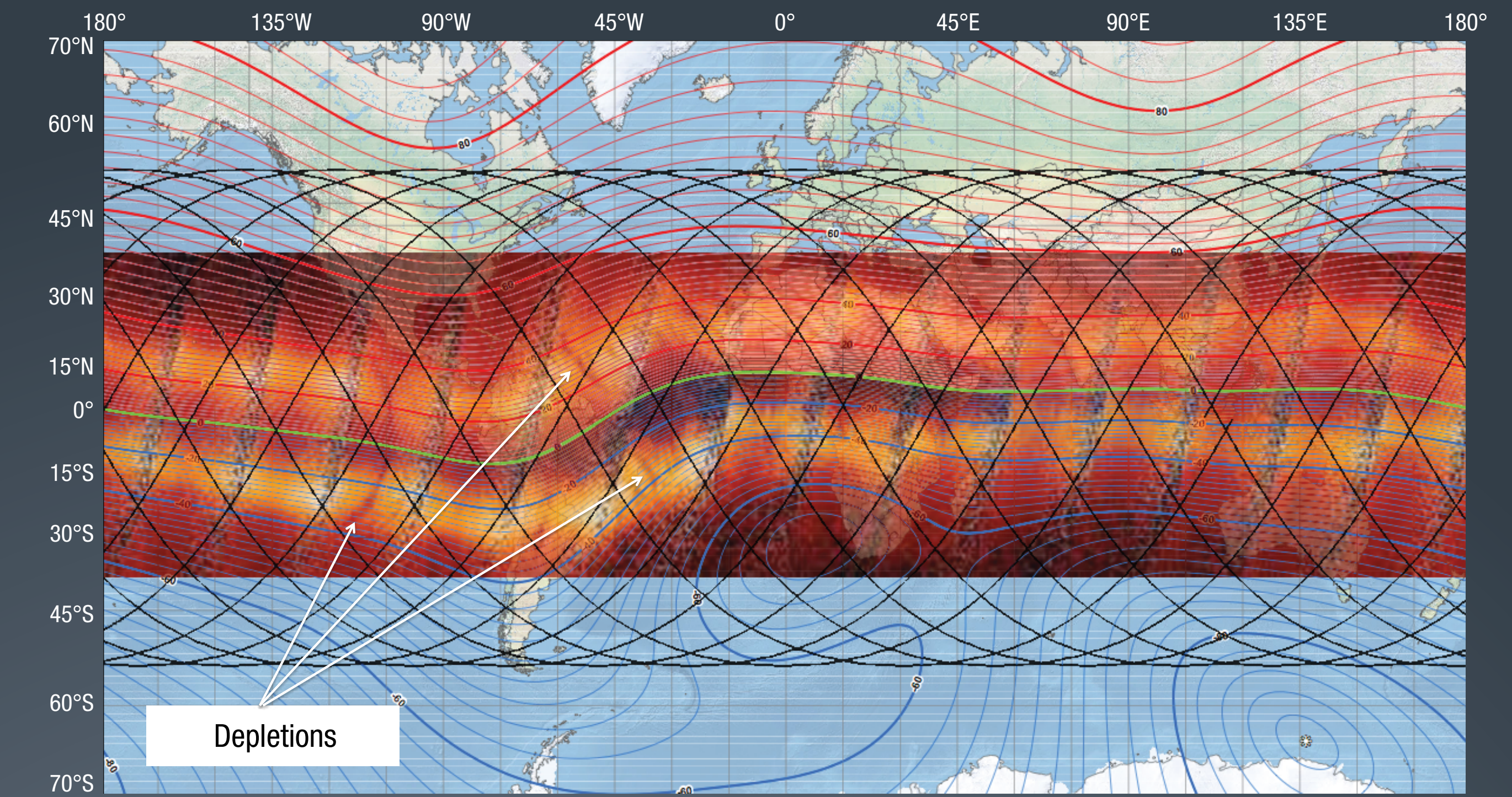
## Science Traceability Matrix

The Scintillation Prediction Observations Research Task (SPORT)		Instrumentation	Spacecraft
Observational Approach	Science Measurement Requirements	Instrument Approach	Space Systems Requirements
1. What is the state of the ionosphere that gives rise to the growth of plasma irregularities that extend into and above the F-peak?			
Observations in the 17:00 to 1:00 LY sector over –30° to 30° latitude	<b>Plasma Density Profile</b> <ol style="list-style-type: none"><li>1. 140 to 450 km alt</li><li>2. 10<sup>4</sup> to 10<sup>7</sup> p/cm<sup>3</sup> range</li><li>3. 20% p/cm<sup>3</sup> accuracy</li><li>4. 1000 km along track sampling</li></ol> <b>Ion Drifts (EarthReference Frame)</b> <ol style="list-style-type: none"><li>1. ±800 m/s Range</li><li>2. 20 m/s precision &amp; accuracy</li><li>3. 10 km along track sampling</li></ol>	<b>GPS Occultation</b> Observe GPS satellite occultation along and to the sides of the orbit plane to obtain line of site TEC  <b>Ion Velocity Meter</b> Observe vertical ion drifts by angle of arrival of heavy ions at detector	<b>Satellite Orbit</b> <ol style="list-style-type: none"><li>1. ≥1 year mission life</li><li>2. 40° to 55° inclination</li><li>3. 350 to 450 km altitude</li><li>4. ±10 km eccentricity</li></ol> <b>Spacecraft</b> <ol style="list-style-type: none"><li>1. ±15° yer mission life</li><li>2. ≤1 km position knowldge</li><li>3. ≤10 ms timing</li></ol>
2. How do plasma irregularities evolve to impact the appearance of radio scintillation at different frequencies?			
Observations in the 22:00 to 2:00 LT sector over over –30° to 30° latitude	<b>E-Field (Earth Reference Frame)</b> <ol style="list-style-type: none"><li>1. ±45 mV/m range</li><li>2. 1.1 mV/m precision &amp; accuracy</li><li>3. 1 km along track sampling</li><li>4. 10 km – 200 m along track waves</li></ol> <b>Plasma Density</b> <ol style="list-style-type: none"><li>1. 10<sup>3</sup> to 10<sup>7</sup> p/cm<sup>3</sup> range</li><li>2. 10<sup>3</sup> p/cm<sup>3</sup> precision &amp; accuracy</li><li>3. 1 km along track sampling</li><li>4. 10 km – 200 m along track waves</li></ol> <b>B-field</b> <ol style="list-style-type: none"><li>1. ± 56,000 nT range</li><li>2. ±100 nT precision and accuracy</li><li>3. 1 km along track sampling</li></ol>	<b>E-Field Double Probe</b> Observe probe floating potential for AC E-fields from irregularity <b>GPS Occultation</b> S4 scintillation index <b>Langmuir/Impedance</b> Observe DC and AC probe response for relative and absolute electron density and observe irregularities <b>Three Axis Magnetometer</b> Support VxB computation for ion velocity and E-Field measurements	<b>Spacecraft Mechanisms</b> <ol style="list-style-type: none"><li>1. ≥0.6 m tip-to-tip booms</li></ol> <b>Attitude (Post Flight Knowledge)</b> <ol style="list-style-type: none"><li>1. ≤0.02° 1σ-uncertainty</li></ol>
Observations of irregularities in electron density and E-field power spectral density in slope from 200 km to 200 m			

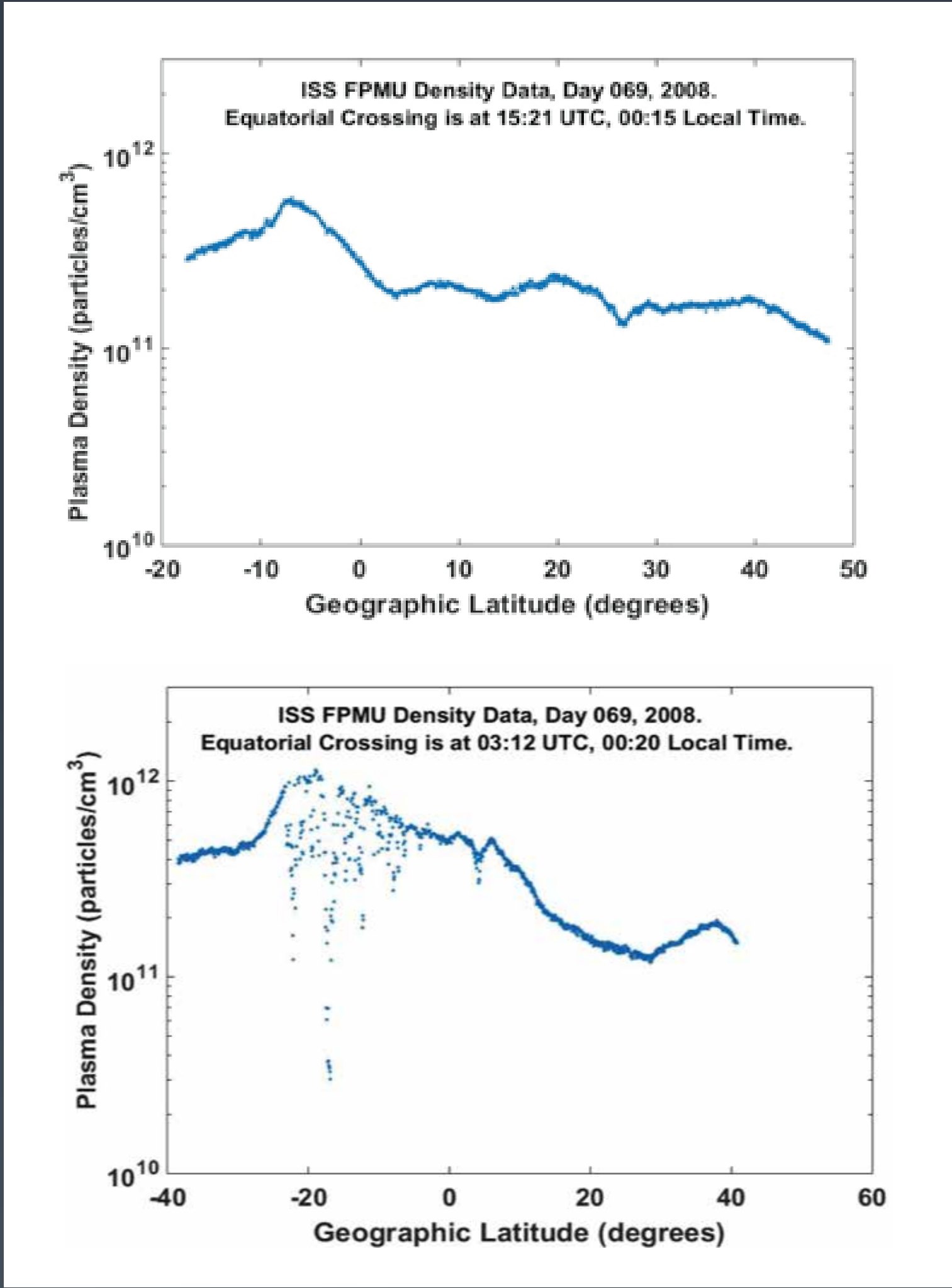
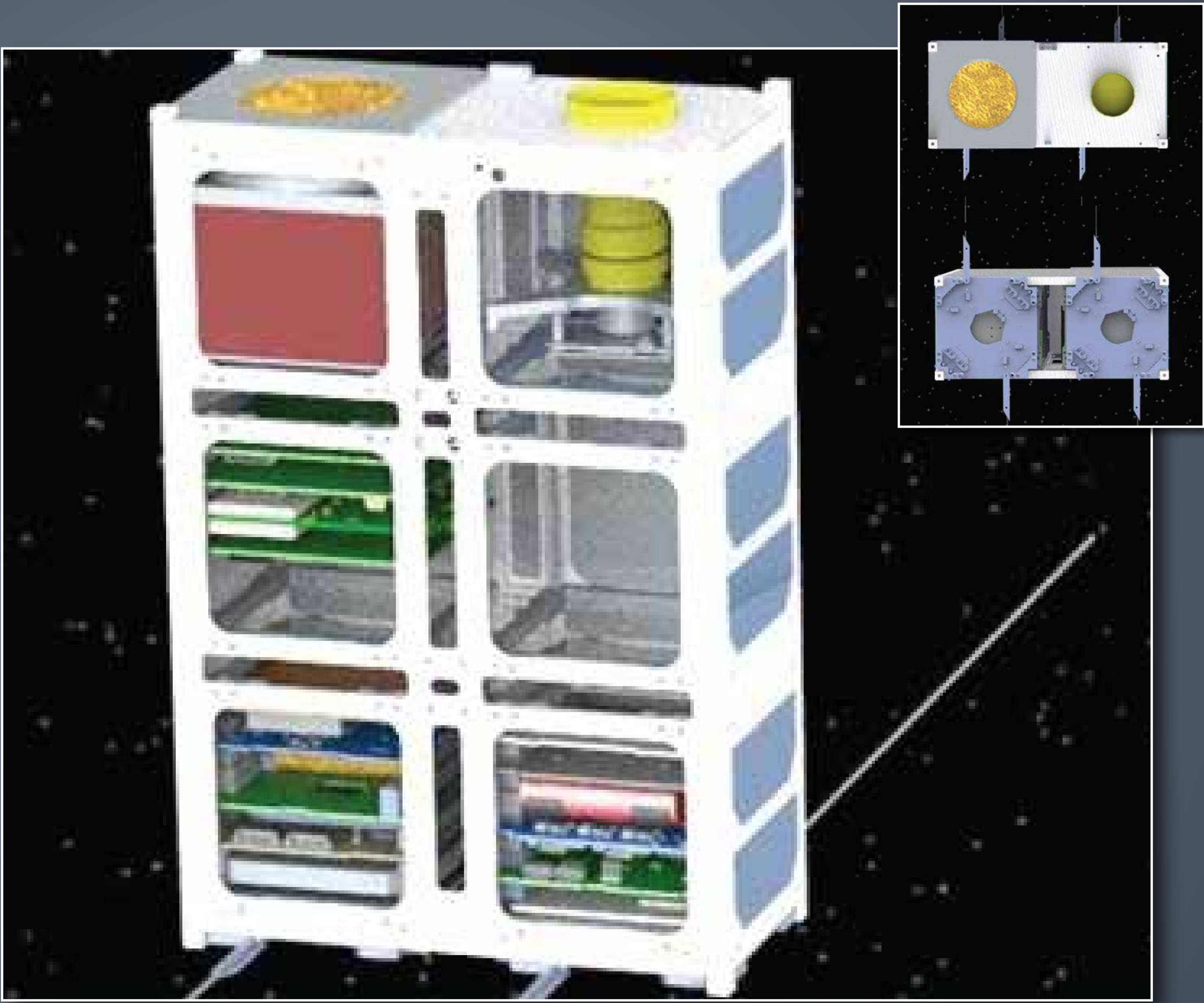
Table 1. Expected Instrument Performance and Requirements

Parameter	Ion Velocity Meter	GPS Occultation	Electric Field Probe	Langmuir Probe	Impedance Probe	Magnetometer
Scientific Requirement	V <sub>i</sub> : ±800 m/s, 20 m/s ΔN <sub>i</sub> : 10 <sup>4</sup> to 10 <sup>7</sup> cm <sup>-3</sup>	N <sub>e</sub> -Profile: 10 <sup>4</sup> to 10 <sup>7</sup> cm <sup>-3</sup> S4 0.2 to 1.2	0.1 to ±45 mV/m	ΔN <sub>e</sub> : 10 <sup>3</sup> to 10 <sup>7</sup> cm <sup>-3</sup> ΔN <sub>i</sub> : 10 <sup>3</sup> to 10 <sup>7</sup> cm <sup>-3</sup>	N <sub>e</sub> : 10 <sup>3</sup> to 10 <sup>7</sup> cm <sup>-3</sup>	± 56,000 nT, 100 nT
Instrument Performance	V <sub>i</sub> : ±1000 m/s, 15 m/s ΔN <sub>i</sub> : 10 <sup>2</sup> to 10 <sup>7</sup> cm <sup>-3</sup> , 5% T <sub>e</sub> : 250 to 5000 K C <sub>i</sub> : 0–100%, 1–40 amu  DC to 2 Hz	Scintillations (S4) Slant TEC: 3 to 200 units Ne-Profile: 10 <sup>3</sup> to 10 <sup>7</sup> cm <sup>-3</sup> S4 0.1 to 1.5 σ: 0.1 to 20 rads 50 Hz	0.1 to 500 mV/m, 1% V <sub>f</sub> (derived): 20 m/s  DC-40 Hz 16 spectrometer ch. 20 Hz to 15 kHz	ΔN <sub>e</sub> : 10 to 10 <sup>7</sup> cm <sup>-3</sup> , 5% ΔN <sub>i</sub> : 10 <sup>3</sup> to 10 <sup>3</sup> cm <sup>-3</sup> , 5% T <sub>e</sub> : 200 to 5000 K V <sub>f</sub> : ±10 mV to ± 12 V V <sub>p</sub> : ±10 mV to ± 12 V DC-40 Hz, 25 s/sweep 16 spectrometer ch. 20 Hz to 15 kHz	N <sub>e</sub> : 10 to 10 <sup>7</sup> cm <sup>-3</sup> , 1%  DC-40 Hz, 25 s/sweep	± 64,000 nT, 10 nT  DC-40 Hz
Mechanism	8 cm aperture	7.6 x 7.6 x 0.5 cm patch antenna	Two 30 cm booms	0.3 x 30 cm boom	30 cm boom	25 cm boom
Attitude Control	15° pointing control	15° pointing control	15° pointing control	15° pointing control	15° pointing control	NA
Attitude knowledge post processed req.	0.02°	2°	0.02°	10°	10°	2° pointing
Field of View	30°	160°	180°	180°	180°	180°
Peak Power	0.3 W	1.5 W	0.15 W	0.15 W	0.4 W	0.45 W
Volume	1.0U Cube  9 × 9 × 10 cm	~0.15U Cube  1.5 × 9 × 9 cm	~0.1U Cube (Shared with LP)  0.75 × 9 × 9 cm	~0.1U Cube (Shared with E-Field)  0.75 × 9 × 9 cm	~0.1U Cube  0.75 × 9 × 9 cm	~0.5U Cube  5 × 9 × 9 cm
Mass	< 1000 g	< 200 g	< 80 g (shared)	< 80g (shared)	< 160 g	< 150 g
Data Rate	2.0 kbps	1.0 kbps Day; 15 kbps Night	1.4 kbps	2.0 kbps	1 kbps	2.8 kbps
Horizontal Cell Size	100 km	500 km	200 m; 20 m spectrometer	200 m; 20 m spectrometer	190 km	10 km
Vertical Cell Size	NA	30 km	NA	NA	NA	NA

V<sub>i</sub> – ion drift velocities; ΔN<sub>i</sub> – relative ion density; ΔN<sub>e</sub> – relative electron density; T<sub>e</sub> – electron temperature; T<sub>i</sub> – ion temperature; V<sub>f</sub> – floating potential; V<sub>p</sub> – plasma potential; N<sub>e</sub> – electron density; B – Magnetic Field; TEC – total electron content; C<sub>i</sub> – ion composition; DC – 1D DC Electric Field; S4 – RF signal amplitude index, σ – RF signal phase index



UV Airglow images from TIMED GUVI clearly showing the equatorial anomaly with embedded depletions that have penetrated through the F peak. Green, Red and Blue traces show the magnetic equator and positive and negative dip angles. SPORT 52° inclination ground tracks are superimposed as black traces.



Total plasma density measurements from the ISS showing signatures across the equatorial anomaly near local midnight. The top panel shows no ionospheric irregularities while the bottom panel shows the presence of depletions and irregularities.

SPORT is science mission on a CubeSat that will advance understanding and improve predictions of scintillatio occurance that impact GPS signals and radio communications. This is science of Space Weather.

SPORT is an international partnership with NASA, the Brazilian space agency research center (INPE), and the Technical Aeronautics Institute under the Brazilian Air Force Command Department (DCTA/ITA).

